## 2012 Project X Physics Study (PXPS12)

14-23 June 2012 Fermi National Accelerator Laboratory
US/Central timezone

# Review of future short baseline accelerator experiments

M. Shaevitz - Columbia University

## Hints for High ∆m<sup>2</sup>~1 eV<sup>2</sup> Oscillation ⇒ Sterile Neutrinos? or Something Else?

#### Positive indications:

	•				
	Anomaly	Type	Channel	Significance	
$\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{e}$	LSND	DAR	$\overline{ u}$ CC	3.8 σ	
$\nu_{\mu} \rightarrow \nu_{e}$	MiniBooNE	SBL accelerator	$\nu$ CC	3.0 σ	New MiniBooNE
$\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{e}$	MiniBooNE	SBL accelerator	$\overline{ u}$ CC	$1.7\sigma$	Combined $v + \overline{v}$
$\nu_e \rightarrow \nu_e$	Gallium/Sage	Source - e capture	$\nu$ CC	$2.7\sigma$	Now 3.8 σ
$\overline{\nu}_{e} \rightarrow \overline{\nu}_{e}$	Reactor	Beta-decay	$\overline{ u}$	$3.0\sigma$	

#### Negative indications:

- CDHS and MiniBooNE restrictions on  $\nu_\mu$  disappearance
- MiniBooNE restrictions on  $\overline{\nu}_{\mu}$  disappearance
- MINOS restrictions on  $\nu_\mu\!\!\to\nu_s$
- Karmen restrictions on  $\overline{\nu}_{\mu} \!\! \to \overline{\nu}_{e}$
- Other negative results

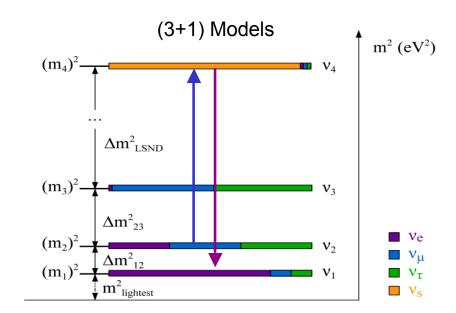


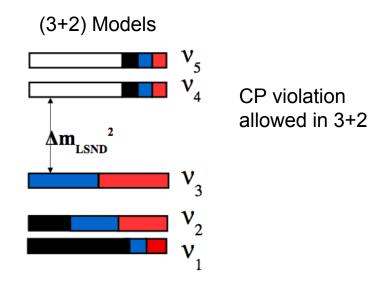
#### **Phenomenology of Oscillations with Sterile Neutrinos**

• In sterile neutrino (3+1) models, appearance comes from oscillation through  $v_s$ 

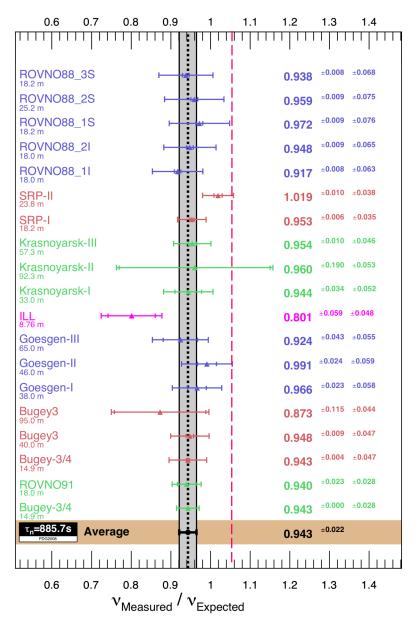
$$- v_{\mu} \rightarrow v_{e} = (v_{\mu} \rightarrow v_{s}) + (v_{s} \rightarrow v_{e})$$

- (3+1) models require  $\nu_{\mu}$  and  $\nu_{e}$  disappearance oscillations
  - $\nu_{\mu} \rightarrow \nu_{s}$  and  $\nu_{e} \rightarrow \nu_{s}$
  - Constraints from disappearance restrict application of (3+1) fits
- Current measurements of appearance and disappearance are not very compatible with (3+1) models ⇒ (3+2) models
  - If  $\nu_{\mu} \rightarrow \nu_{e}$  and  $\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{e}$  are different then (3+2) models can have CP violation
  - Still tension between appearance and disappearance





#### **Reactor Antineutrino Anomaly**



#### New Reactor antineutrino Spectra

- Net 3% upward shift in energy-averaged fluxes
- Phys. Rev. C83, 054615, 2011

## Recent re-analysis of 19 reactor neutrino results

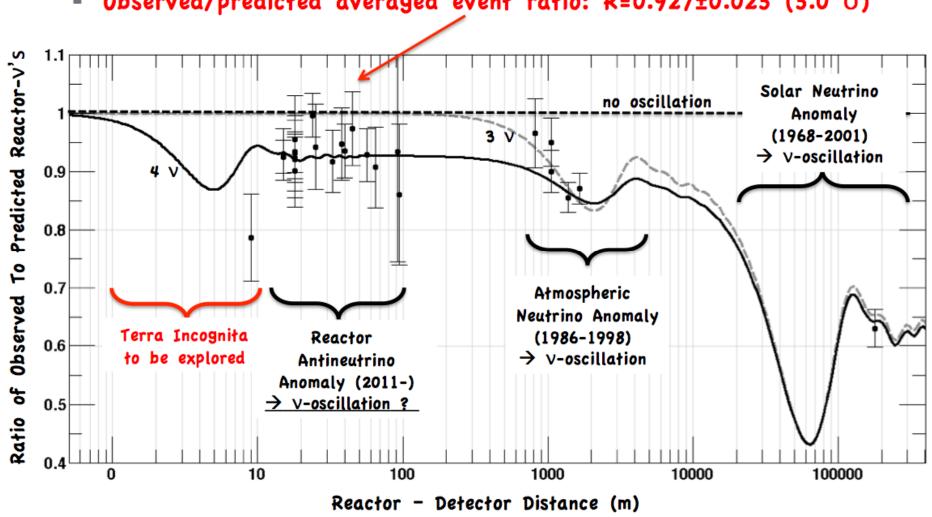
- Neutron life time correction & Offequilibrium effects
- Phys. Rev. D83, 073006, 2011
- Obs/Pred =  $0.927\pm0.023$  (3  $\sigma$ )

#### •At least three alternatives:

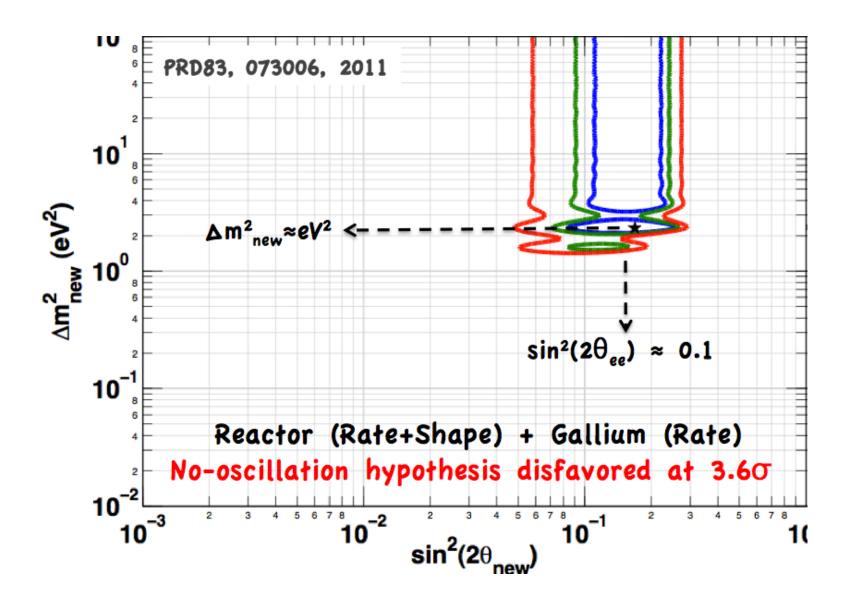
- Wrong prediction of v-spectra ?
- Bias in all experiments ?
- New physics at short baselines: Mixing with 4<sup>th</sup> v-state

#### **Reactor Antineutrino Anomaly**

Observed/predicted averaged event ratio: R=0.927±0.023 (3.0  $\sigma$ )

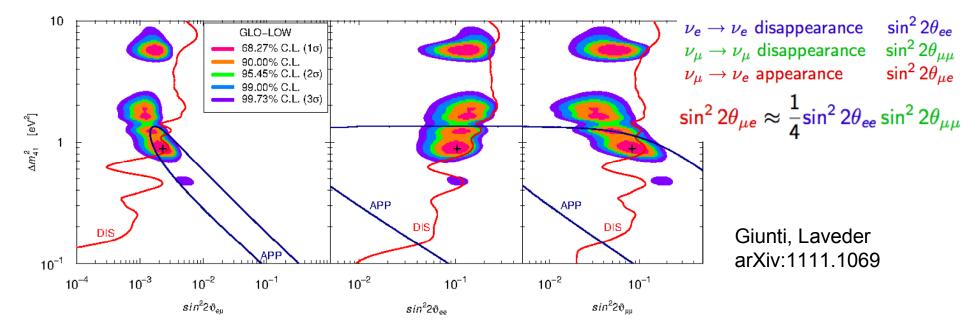


#### **Combined Gallium and Reactor Allowed Region (3+1)**



#### Global Fits to Appearance and Disappearance Results

- In 3+1 models, hard to reconcile  $v_e/v_e$  appearance/disappearance with  $\overline{v_u}/v_u$  disappearance
  - Compatibility among data sets for 3+1 fits less than 1%



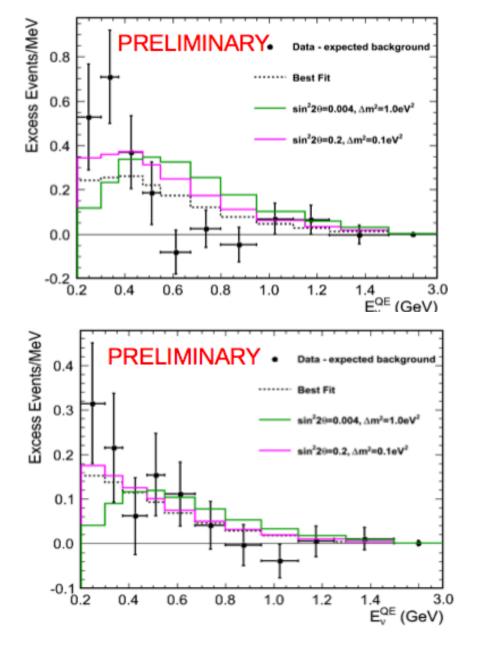
• 3+2 models better since there can be CP violating interference appearance:

$$P_{\nu_{\mu} \to \nu_{e}} = 4 |U_{e4}|^{2} |U_{\mu 4}|^{2} \sin^{2} \phi_{41} + 4 |U_{e5}|^{2} |U_{\mu 5}|^{2} \sin^{2} \phi_{51} + 8 |U_{e4}U_{\mu 4}U_{e5}U_{\mu 5}| \sin \phi_{41} \sin \phi_{51} \cos(\phi_{54} - \delta)$$

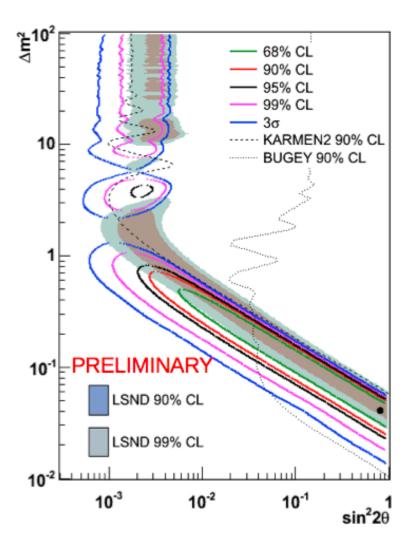
disappearance:

$$P_{\nu_{\alpha} \to \nu_{\alpha}} \approx 1 - 4 \sum_{i=4,5} |U_{\alpha i}|^2 \sin^2 \phi_{i1} - 4 |U_{\alpha 4}|^2 |U_{\alpha 5}|^2 \sin^2 \phi_{54}$$

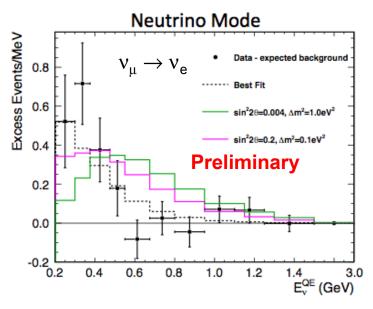
## New MiniBooNE $\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{e}$

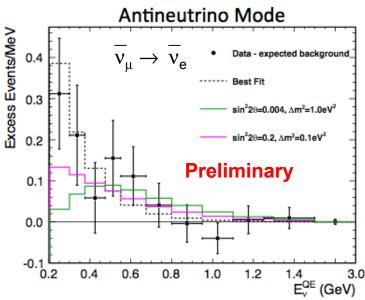


•  $v_{\mu} \rightarrow v_{e}$  and  $\overline{v}_{\mu} \rightarrow \overline{v}_{e}$  becoming more compatible with a common oscillation hypothesis and with the LSND result

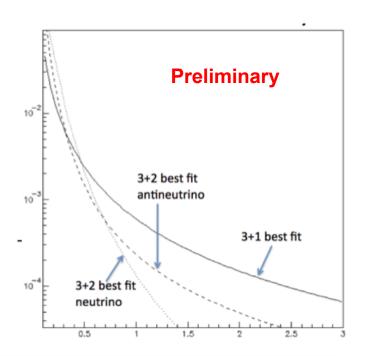


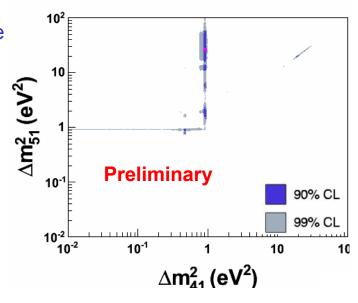
### Preliminary (3+2) Fits to New MiniBooNE $v_e$ / $\overline{v_e}$ Appearance





- Two high mass scales plus CP violation effects can possibly explain v<sub>e</sub> vs v<sub>e</sub> appearance
- Still some tension with disappearance results.





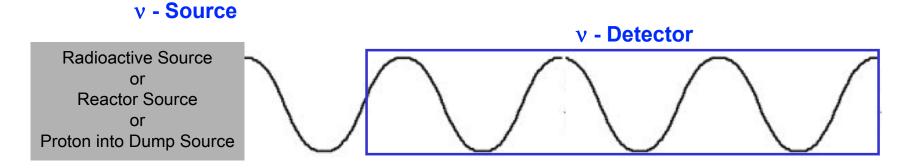
#### **Many Ideas for Future Experiments**

- Establishing the existence of sterile neutrinos would be a major result for particle physics
- Need definitive experiments
  - Significance at the  $> 5\sigma$  level
  - Observation of oscillatory behavior (L and/or E dependence) within a detector or between multiple detectors
  - Oscillation signal clearly separated from backgrounds
- Need to make both appearance and disappearance oscillation searches for neutrinos and for antineutrinos
  - Needed to prove the consistency with sterile neutrino (3+1) and (3+2) models
- Very active area for the field with many proposals and ideas
  - "Light Sterile Neutrinos: A White Paper" (arXiv:1204.5379) put together by a group of over 170 experimentalists and theorists.
  - Many workshops investigating opportunities and possibilities

## **Future Experimental Oscillation Proposals**

Type of Exp	App/Disapp	Osc Channel	Experiments
Reactor Source	Disapp	$\overline{\nu}_{e}  ightarrow \overline{\nu}_{e}$	Nucifer, Stereo, SCRAMM, NIST, Neutrino4, DANSS
Radioactive Sources	Disapp	$ \begin{array}{c} \overline{\nu}_{e} \rightarrow \overline{\nu}_{e} \\ (\nu_{e} \rightarrow \nu_{e}) \end{array} $	Baksan, LENS, Borexino, SNO+, Richochet, CeLAND, Daya-Bay
Isotope Source	Disapp	$\overline{\nu}_{e} \rightarrow \overline{\nu}_{e}$	IsoDAR
Pion / Kaon Decay- at-Rest Source	Appearance & Disapp	$ \begin{array}{c} \overline{\nu}_{\mu} \to \overline{\nu}_{e} \\ \nu_{e} \to \nu_{e} \end{array} $	OscSNS, CLEAR, DAEδALUS, KDAR
Accelerator $\stackrel{(-)}{\nu}$ using Pion Decay-in-Flight	Appearance & Disapp	$ \begin{vmatrix} \nu_{\mu} \rightarrow \nu_{e} \ , \ \overline{\nu}_{\mu} \rightarrow \overline{\nu}_{e} \\ \nu_{\mu} \rightarrow \nu_{\mu} \ , \ \nu_{e} \rightarrow \nu_{e} \end{vmatrix} $	MINOS+, MicroBooNE, LAr1kton+MicroBooNE, CERN SPS
Low-Energy v-Factory	Appearance & Disapp	$\begin{array}{c} \nu_{e} \rightarrow \nu_{\mu} \; , \; \overline{\nu}_{e} \rightarrow \overline{\nu}_{\mu} \\ \nu_{\mu} \rightarrow \nu_{\mu} \; , \; \nu_{e} \rightarrow \; \nu_{e} \end{array}$	νSTORM at Fermilab

#### **Very-short Baseline Oscillation Experiments**



 $1/L^2$  flux rate modulated by  $\text{Prob}_{osc} = \sin^2 2\theta \cdot \sin^2 \left( \Delta m^2 L / E \right)$ 

- Can observe oscillatory behavior within the detector if neutrino source has small extent.
  - Look for a change in event rate as a function of position and energy within the detector
  - Bin observed events in L/E (corrected for the 1/L<sup>2</sup>) to search for oscillations
- Backgrounds produce fake events that do not show the oscillation L/E behavior and are easily separated from signal

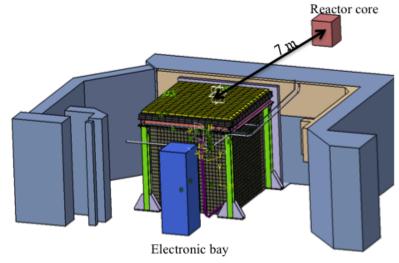
Very-Short Baseline Reactor Experiments (  $\overline{\nu}_e$  Disappearance )

#### **Very-Short Baseline Reactor Experiments**

Experiment	Reactor	Baseline	Status
Nucifer (Saclay)	Osiris 70MW	7	Taking Data
Stereo (Genoble)	ILL 50 MW	10	Proposal
SCRAMM (CA)	San-Onofre 3 GW	24	Proposal
NIST (US)	NCNR 20 MW	4-11	Proposal
NEUTRINO4	SM3 100 MW	6-12	Proposal
SCRAMM (Idaho)	ATR 150 MW	12	Proposal
DANSS (Russia)	KNPP 3 GW	14	Fabrication

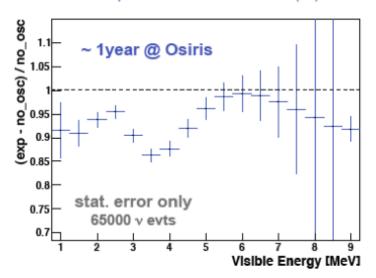
#### **NUCIFER Reactor Experiment**

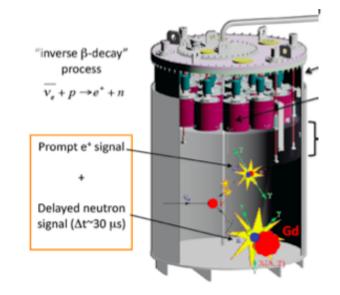
Osiris Research Reactor: Core Size: 57x57x60 cm 1.2m x 0.7m detector , 7m distance from core

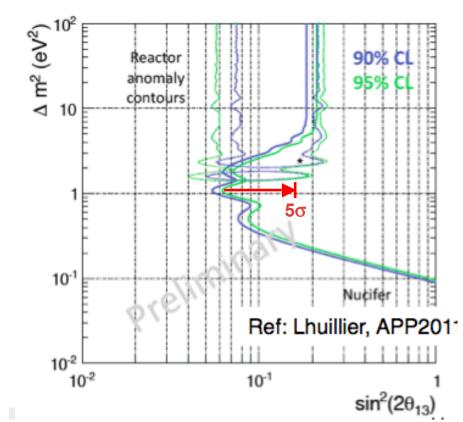


#### Expected E spectrum deformation

with anomaly best fit:  $\Delta m^2 = 2.4 \text{ eV}^2 \& \sin^2(2\theta) = 0.15$ 







## Radioactive $\beta$ -Decay Source Experiments ( $v_e$ or $\overline{v_e}$ Disappearance)

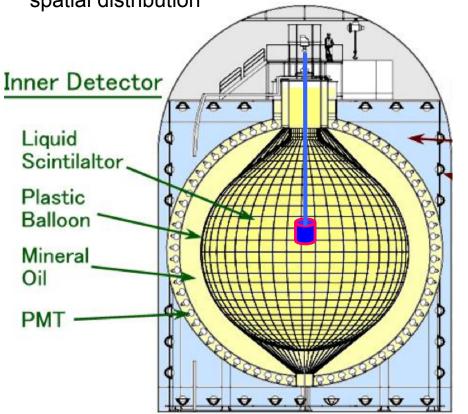
### Radioactive $\beta$ -Decay Source Experiments

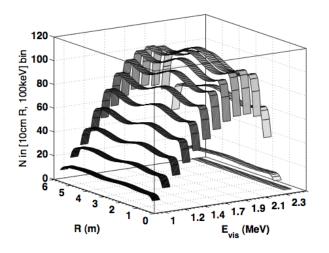
Species	Source	Experiment	Status
V <sub>e</sub>	<sup>51</sup> Cr	Baksan	Proposal
V <sub>e</sub>	<sup>51</sup> Cr	LENS	Proposal
V <sub>e</sub>	<sup>51</sup> Cr	Borexino	Proposal
V <sub>e</sub>	<sup>51</sup> Cr	SNO+	Proposal
V <sub>e</sub>	<sup>37</sup> Ar	Richochet	Proposal
$\overline{v}_{e}$	<sup>144</sup> Ce	Ce-LAND	Proposal
$\overline{v}_{e}$	<sup>144</sup> Ce	Daya-Bay	Proposal

#### Ce-LAND Exp: Using 144Ce kCi Anti-neutrino Source

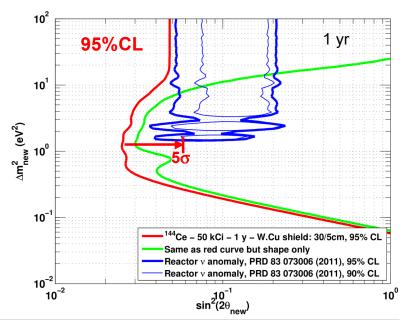
- A 50 kCi anti-v source (10 g of <sup>144</sup>Ce) in the middle of a large LS detector
- Inside a thick 35 cm W-Cu shielding → background free

Energy-dependent oscillating pattern in event spatial distribution





M. Cribier, et al. PRL 107, 201801(2011)

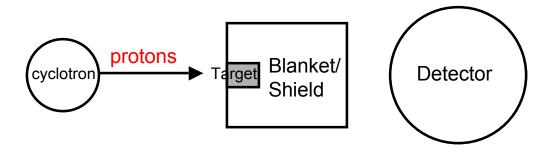


Detectors which could be used for this idea include Kamland, SNO+, or Borexino...

Isotope Decay-at-Rest Neutrino Source (  $\overline{\nu_e}$  Disappearance )

## **IsoDAR** $\overline{v}_e$ **Disappearance Exp** (arXiv:1205.4419)

- High intensity  $\overline{\nu}_e$  source using  $\beta$ -decay at rest of  $^8$ Li isotope  $\Rightarrow$  IsoDAR
- <sup>8</sup>Li produced by high intensity (10ma) proton beam from 60 MeV cyclotron
   ⇒ being developed as prototype injector for DAEδALUS cyclotron system
- Put a cyclotron-isotope source near one of the large (kton size) liquid scintillator/water detectors such as KAMLAND, SNO+, Borexino, Super-K....



- Physics measurements:
  - $\overline{v}_e$  disappearance measurement in the region of the LSND and reactorneutrino anomalies.
  - Measure oscillatory behavior within the detector.

#### **IsoDAR 60 MeV Proton Cyclotron (Under Development)**

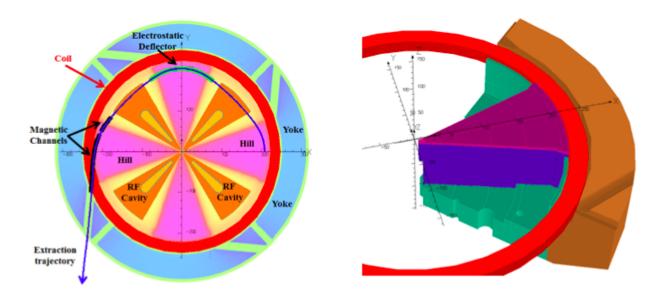
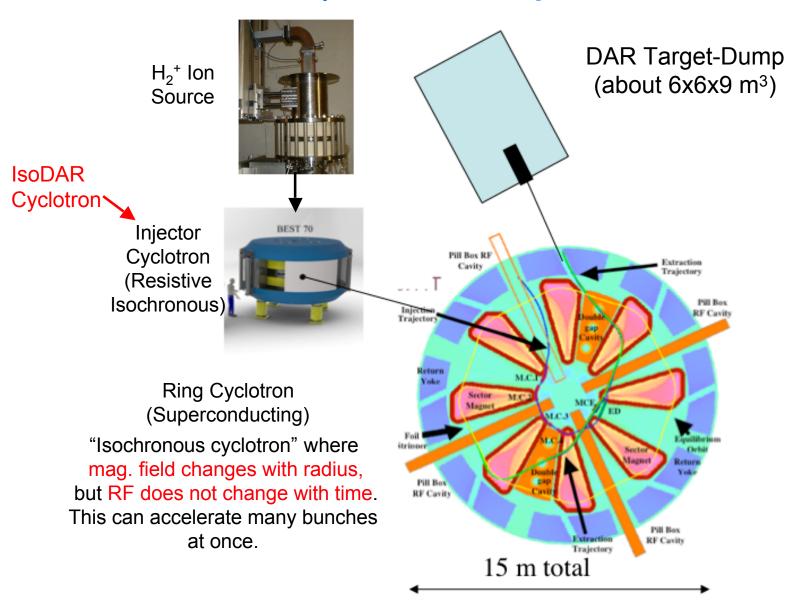


Figure 3: Left: Layout of the injector cyclotron. Pastel colors indicate magnetic field map (pink is highest). The hill/valley structure is apparent. Extraction trajectory for  $H_2^+$  is shown. Right: Illustration of the Opera3D finite element magnetic model showing one quarter of the cyclotron with the pole, the return yoke and the coil.

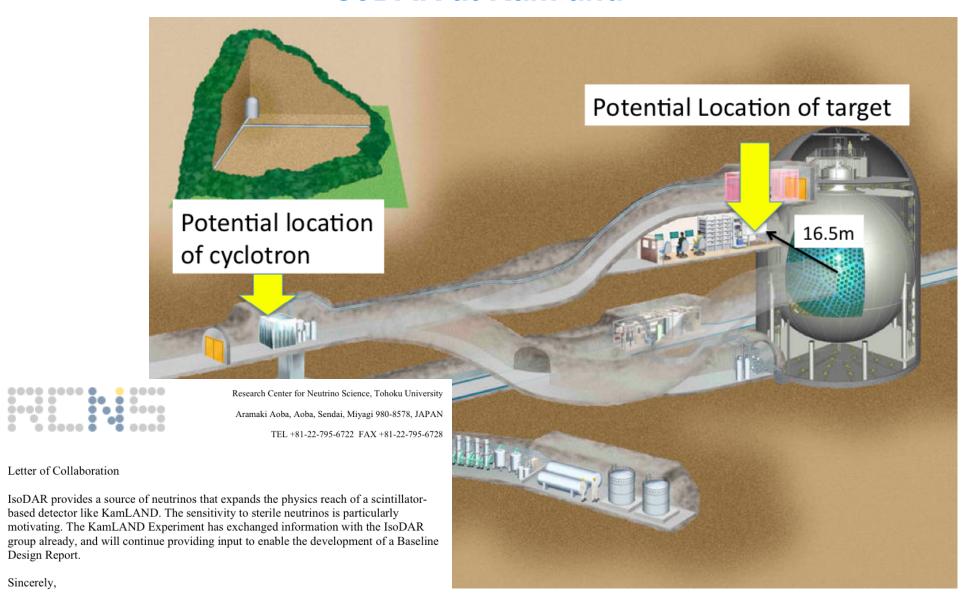
Table 1: Parameters of the DAEδALUS injector cyclotron

	and arrived on of the 1	TILLOTTIL O D ON GOOD	· ogorourore
$E_{max}$	60  MeV/amu	$E_{inj}$	35  keV/amu
$R_{ext}$	1.99  m	$R_{inj}$	55  mm
$\langle B \rangle @ R_{ext}$	1.16 T	$\langle B \rangle @ R_{inj}$	0.97  T
Sectors	4	Hill width	28 - 40 deg
Valley gap	1800  mm	Pole gap	100  mm
Outer Diameter	6.2 m	Full height	$2.7 \mathrm{m}$
Cavities	4	Cavity type	$\lambda/2$ , double gap

## DAEδDALUS 800 MeV Cyclotron System (Under Development)



#### IsoDAR at Kamland



Kunio Iroul

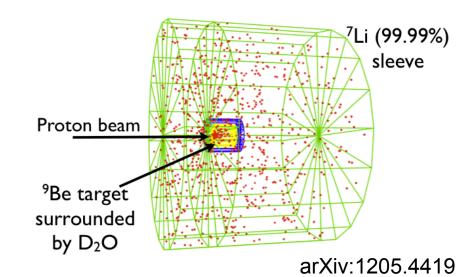
Kunio Inoue Research Center for Neutrino Science, Tohoku University

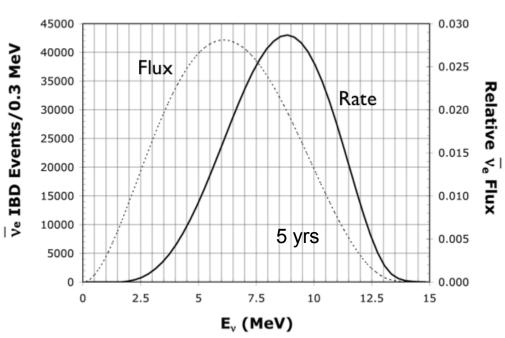
#### **IsoDAR Neutrino Source and Events**

- p (60 MeV) +  ${}^{9}\text{Be} \rightarrow {}^{8}\text{Li} + 2\text{p}$ 
  - plus many neutrons since low binding energy
- n + <sup>7</sup>Li (shielding) → <sup>8</sup>Li

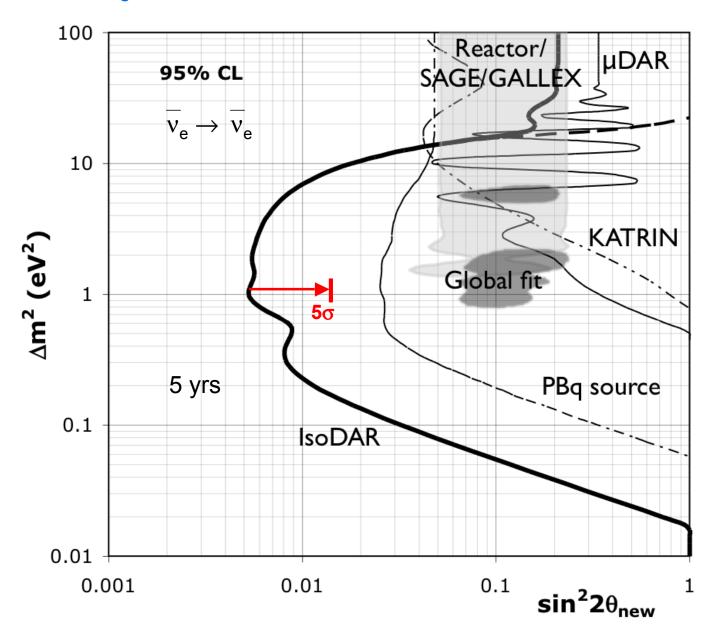
• 
$$^8\text{Li} \rightarrow ^8\text{Be} + e^- + \stackrel{-}{\nu_e}$$

- Mean  $\overline{v}_e$  energy = 6.5 MeV
- $2.6 \times 10^{22} \ \overline{\nu}_{e} \ / \ yr$
- Example detector: Kamland (900 t)
  - Use IBD  $\overline{\nu}_e$  + p  $\rightarrow$  e<sup>+</sup> + n process
  - Detector center 16m from source
  - ~160,000 IBD events / yr
  - 60 MeV protons @ 10ma rate
  - Observe changes in the IBD rate as a function of L/E



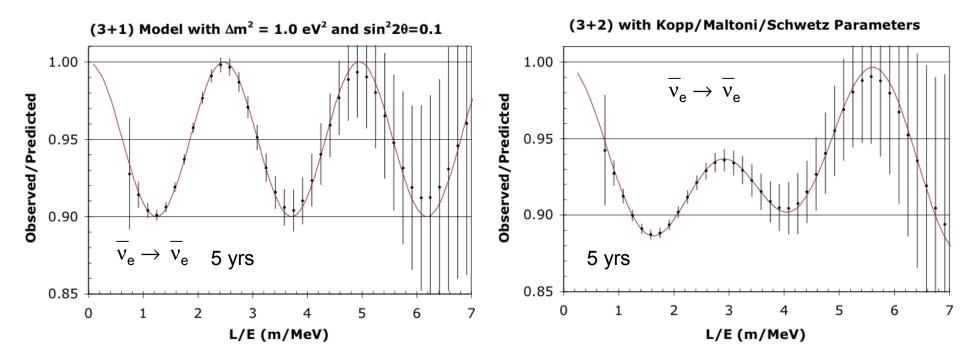


## IsoDAR $\overline{v}_e$ Disappearance Oscillation Sensitivity (3+1)



#### Oscillation L/E Waves in IsoDAR

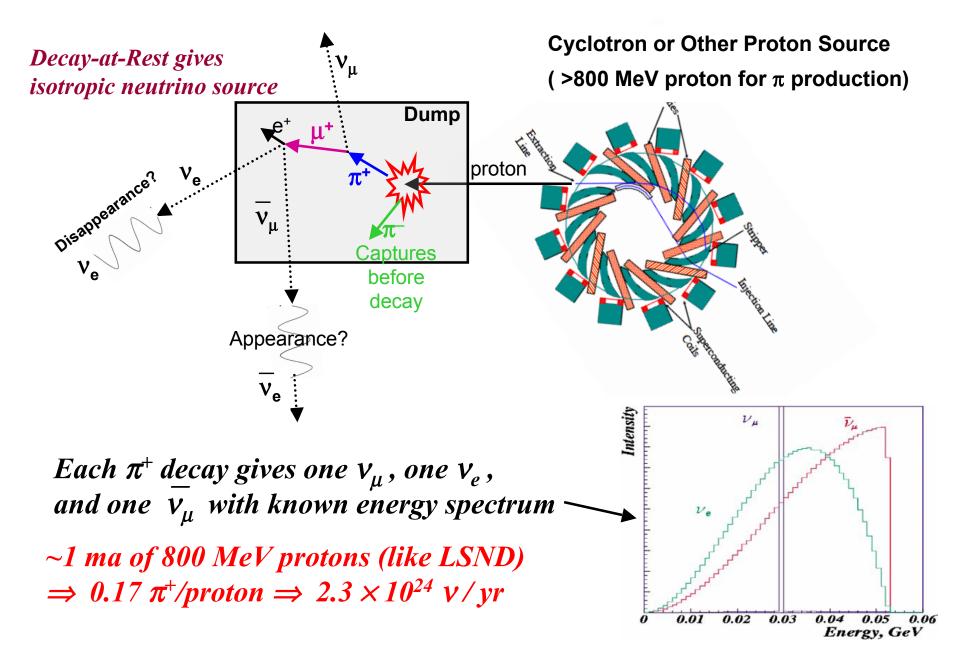
Observed/Predicted event ratio vs L/E including energy and position smearing



IsoDAR's high statistics and good L/E resolution gives good sensitivity to distinguish (3+1) and (3+2) oscillation models

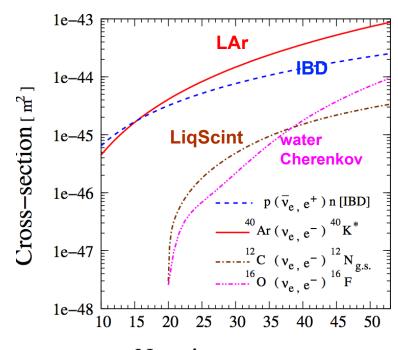
## Pion or Kaon Decay-at-Rest Neutrino Sources

#### **Decay-at-Rest (or Beam Dump) Neutrino Sources**



#### **Short Baseline Osc Exps using DAR Sources**

- Good oscillation sensitivity for DAR  $v / \overline{v}$  -source placed near large detector
  - Neutrino source has small extent (± 25 cm) and can be close (~20m) to detector
  - Energy range 20 to 50 MeV
  - Possible to observe L/E oscillations within the detector
- Detectors: Cherenkov (water or oil), liquid argon, or liquid scintillator
- $v_e \rightarrow v_e$  Disappearance
  - Process: Charged current scattering ( $v_e + N \rightarrow e^- + N$ ')
  - Look for an oscillations in  $\nu_{\text{e}}$  rate with L/E
  - Backgrounds do not have this L/E behavior
- $\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{e}$  Appearance
  - Process: Inverse Beta Decay (IBD)  $v_e + p \rightarrow e^+ + n$
  - Look for oscillation wave in L/E
  - Detector needs to free hydrogen targets and be able to capture the outgoing n
    - ⇒ Only water or liquid scintillator (with Gd better)



Neutrino energy [MeV]

#### **Scintillation Detectors with DAR Neutrino Sources**

0.01

0.009

0.008 0.007

0.006

0.004

0.003

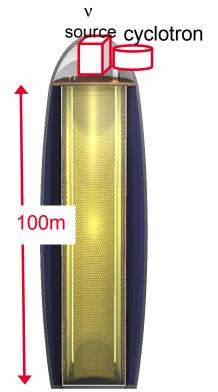
0.002

0.001

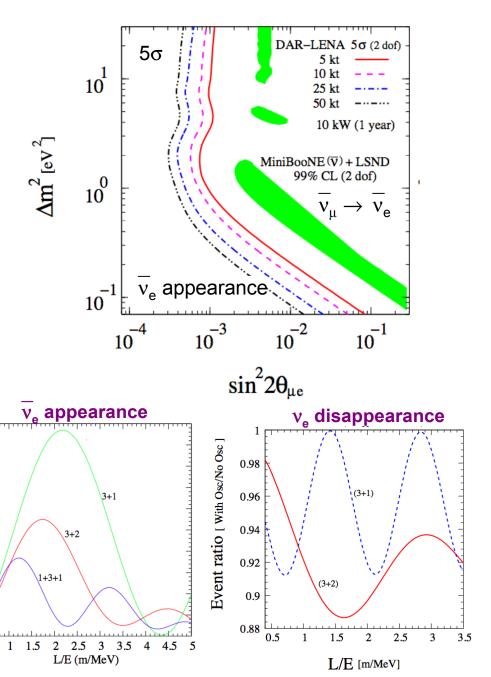
Event ratio (With Osc/Full Transmutation)

## Example: LENA Scintillation Detector (Part of the European LAGUNA Project)

- For 5σ coverage, only need 10 kW source with 5 kton detector
- Deep location (4000 mwe) so minimal cosmic muon backgrounds
- Appearance and Disappearance possible



Agarwalla, Conrad, and MHS: arXiv:1105.4984 (JHEP 1112 (2011) 085)

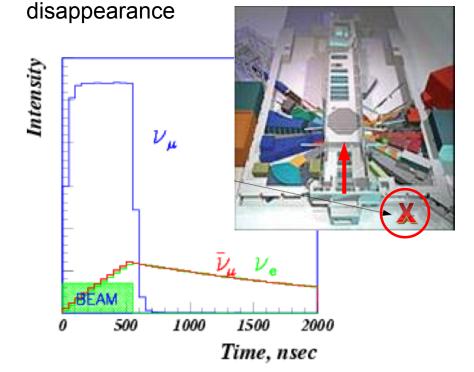


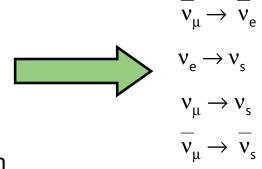
#### OscSNS: DAR Neutrino Source at SNS (ORNL)

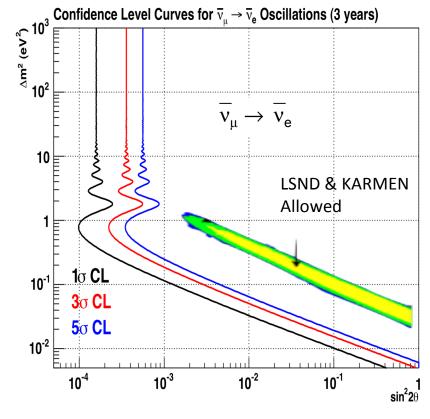
arXiv:0810.3175

- Spallation neutron source at ORNL
- ~1GeV protons on Hg target (1.4MW)
- 6.2% Duty factor reduces backgrounds
- Time structure 695ns <u>p</u>ulses at 20 Hz can separate  $\nu_{\mu}$  from  $\nu_{\mu}$  and  $\nu_{e}$
- 800 ton MiniBooNE style detector 60m from target

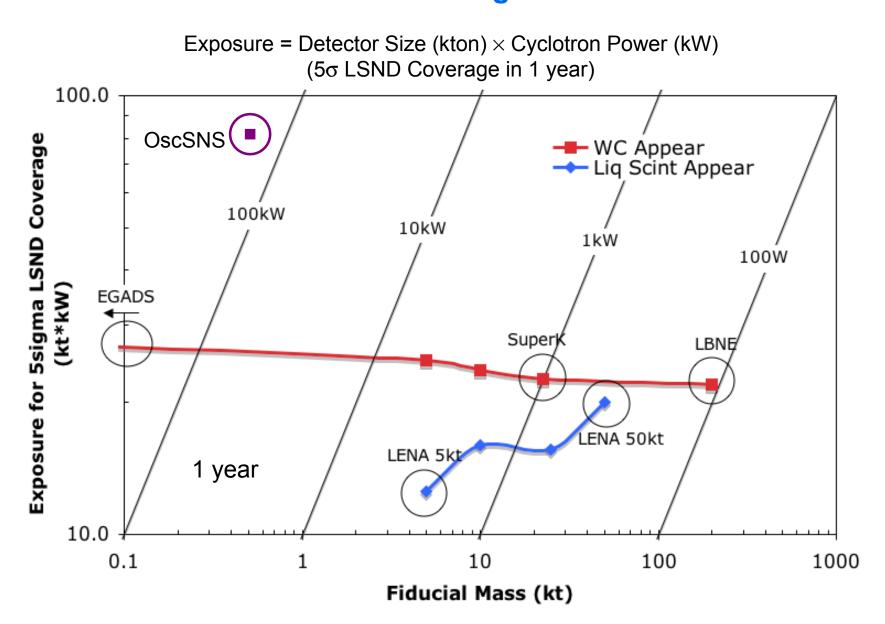
• Can do  $\overline{v}_{e}$  appearance and other types of





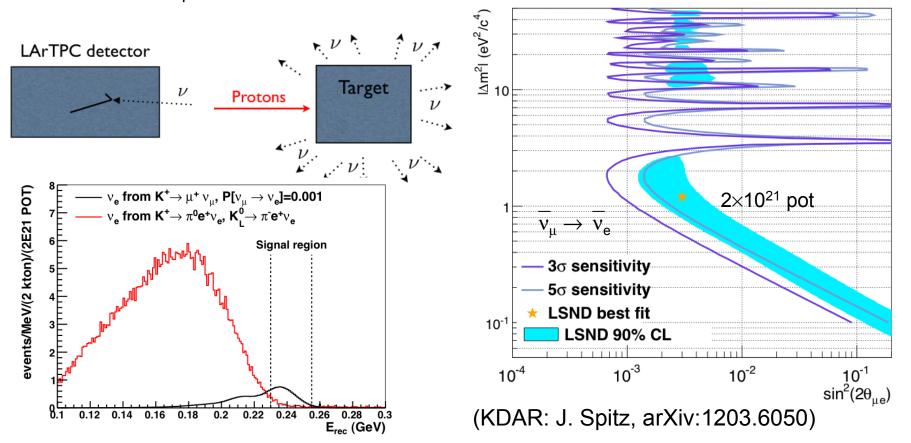


## VSBL $\overline{\nu}_e$ Appearance: Source Power and Detector Size for LSND Coverage at $5\sigma$



#### **Kaon Decay-at-Rest Experiment**

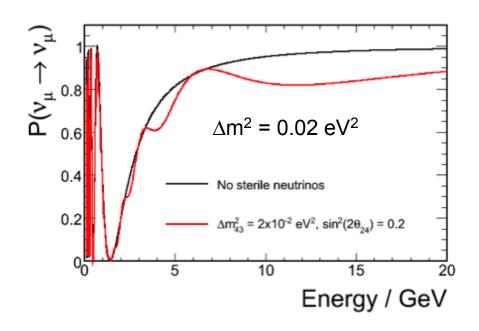
- >10 GeV high-intensity, proton beam into target-dump to produce kaons that stop and decay at rest.
  - Gives a monoenergetic muon neutrinos (235 MeV) from K $\rightarrow \mu + \nu_{\mu}$
- 2 kton LAr detector placed at 160m in backwards direction.
- Look for  $v_{\mu} \rightarrow v_{e}$  oscillations by identifying  $v_{e}$  events at high energy

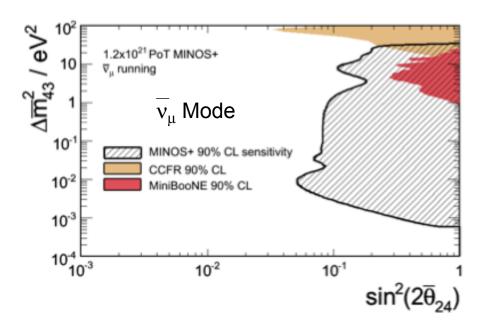


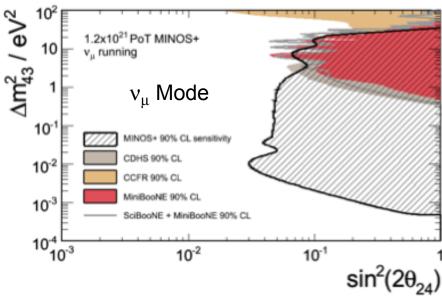
# Accelerator $v_{\mu}$ / $\overline{v_{\mu}}$ Beams using Pion Decay-in-Flight

#### MINOS+ Running (3 yrs) During Nova Era

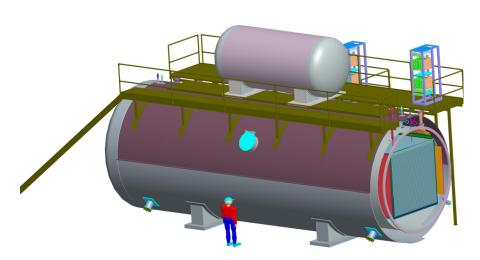
- MINOS+ Sensitivity to sterile neutrinos through neutral current (NC) disappearance between near and far detector
  - If disappearance seen, must be to a sterile neutrino with no NC interactions
- Sensitivity to ∆m² values to below 0.01 eV²



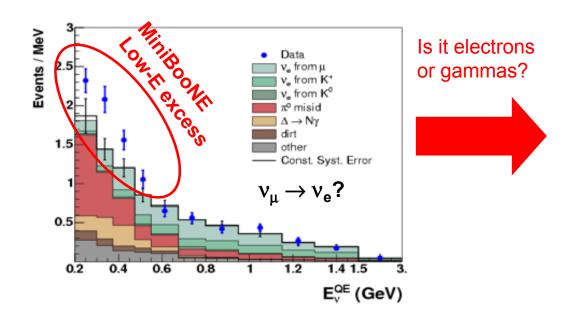


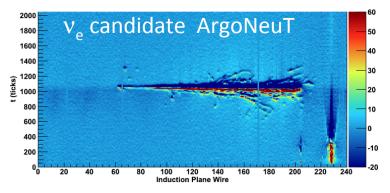


## MicroBooNE Experiment (Under Construction) using Fermilab Booster Neutrino Beamline (BNB)

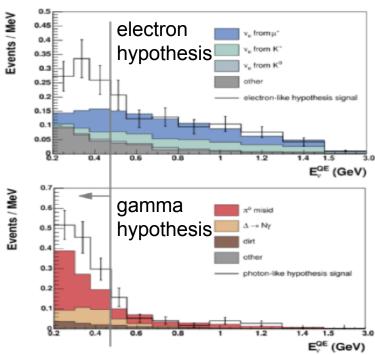


See poster #167 G. Karagiorgi





Use topology and dE/dx to differentiate electrons (signal) from gammas (background) (Indistinguishable in Cerenkov imaging detectors)



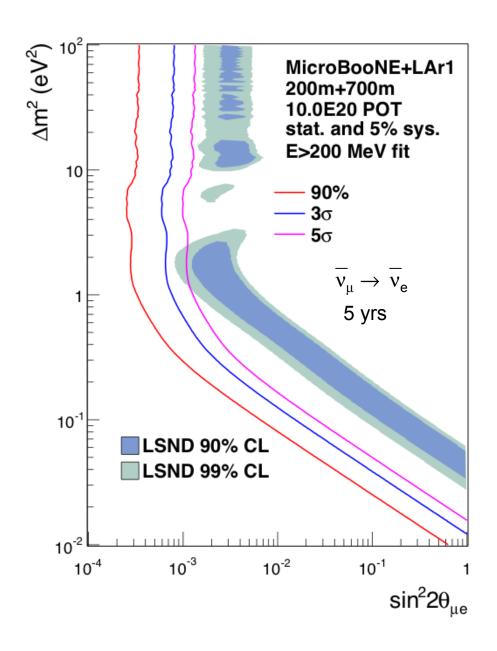
## LAr1kton at Fermilab Booster v Beamline (BNB)

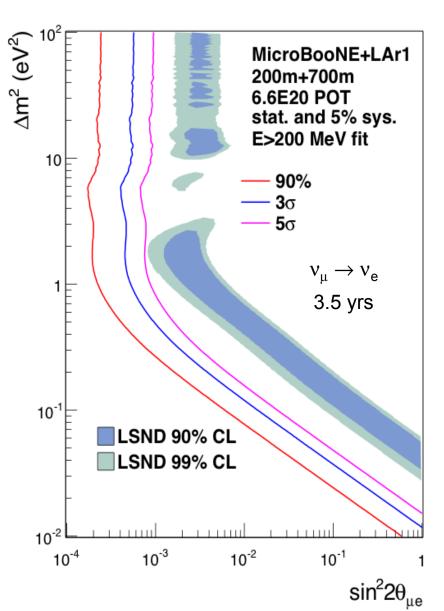
- To directly address LSND  $\overline{\nu}_{\mu} \to \overline{\nu}_{e}$  appearance signal, use multiple detectors in the Fermilab BNB
- Large (1 kton fiducial) LAr detector at 700m plus MicroBooNE at 200m (also maybe MiniBooNE with scintillator at 540 m)

 LAr capabilities significantly reduces gamma and other backgrounds



# **LAr1kton Sensitivity**





# CERN SPS: Two (or Three) Detector Proposal using Liquid Argon and Iron Spectrometers

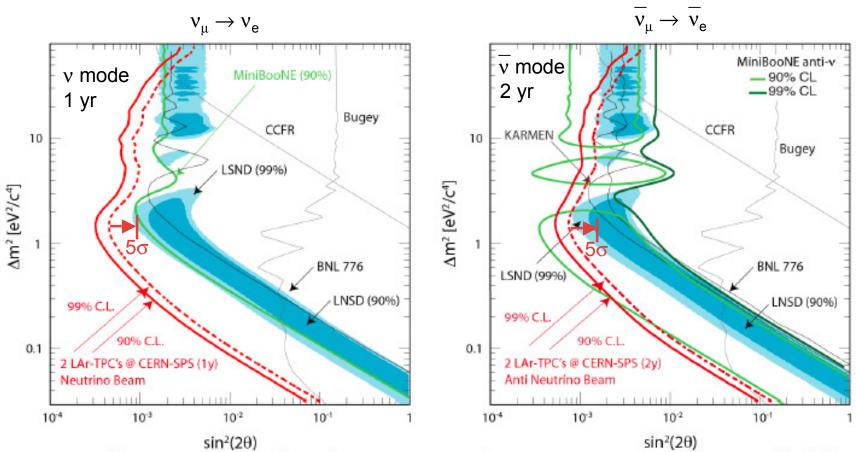
Combined ICARUS and NESSiE Collaborations

# New Neutrino Facility in the CERN North Area



100 GeV primary beam fast extracted from SPS; target station next to TCC2; decay pipe l = 100m,  $\emptyset = 3m$ ; beam dump: 15m of Fe with graphite core, followed by  $\mu$  stations.

# **CERN SPS Appearance Sensitivity**



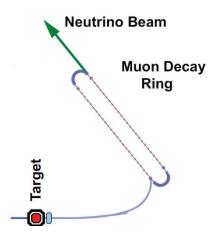
Expected sensitivity for the proposed experiment:  $v_{\mu}$  beam (left) and anti- $v_{\mu}$  (right) for 4.5 10<sup>19</sup> pot (1 year) and 9.0 10<sup>19</sup> pot (2 years) respectively. LSND allowed region is fully explored in both cases.

Also,  $v_{\mu}$  and  $\overline{v_{\mu}}$  disappearance

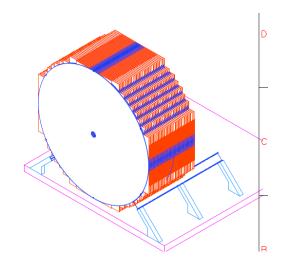
Very Low Energy Neutrino Factory  $v / \overline{v}$  Source

### **Neutrinos from STORed Muons - vSTORM**

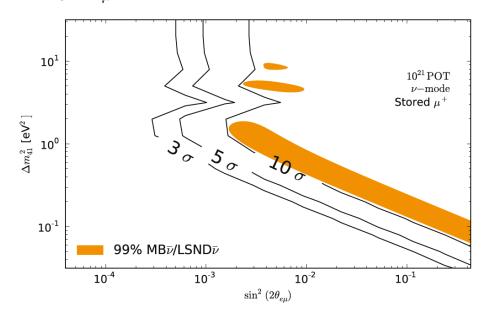
- Simplest implementation of the NF concept
  - 60 GeV protons on solid target (100 kW)
  - Horn capture and  $\pi$  transfer
  - Decay ring
- No new technology is required
  - Little R&D is needed ≈ "Technology" ready



- Performance assumptions:
  - $10^{21} 60 \, \text{GeV/c POT}$
- Yields  $\approx 2X10^{18}$  useful v
- ≈ 2000 m baseline
- 1.3 kT Minos-like detector: SuperB IND
  - Thinner plates
  - 2T B



 $v_e \rightarrow v_u$ : CPT Invariant mode of LSND/MinBooNE



# **Summary and Conclusions**

- Establishing the existence of sterile neutrinos would be a major result for particle physics
- Many proposals and ideas for sterile neutrino searches in the  $\Delta m^2 \sim 1 \text{ eV}^2$  region
  - New experiments have better sensitivity ( $\sim 5\sigma$  level) with capabilities to see oscillatory behavior and reduce backgrounds
- Many different techniques, neutrino sources, and proposals

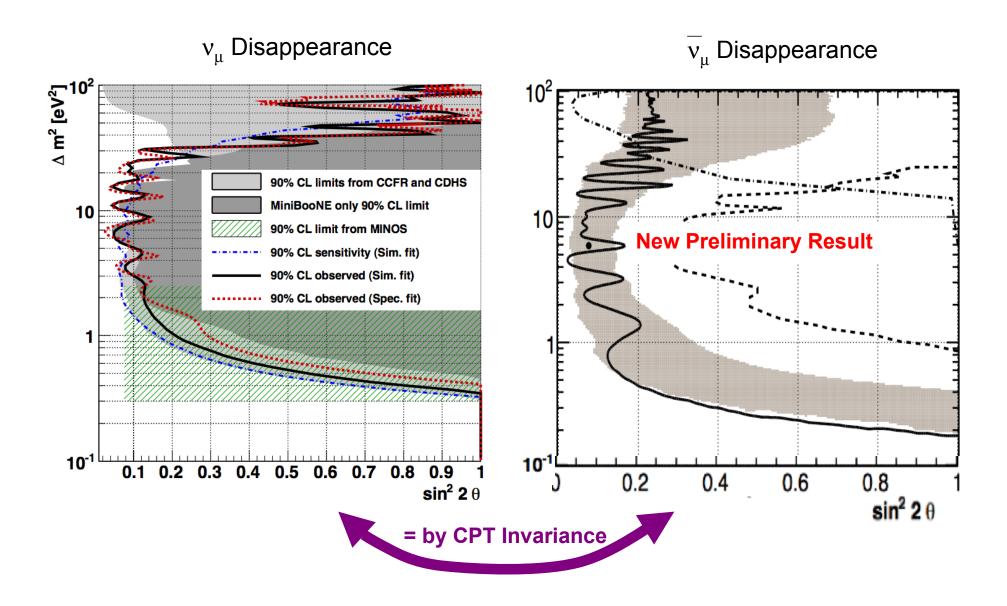
Type of Exp	App/Disapp	Osc Channel	Experiments
Reactor Source	Disapp	$\stackrel{-}{v_e} \rightarrow \stackrel{-}{v_e}$	Nucifer, Stereo, SCRAMM, NIST, Neutrino4, DANSS
Radioactive Sources	Disapp	$ \begin{array}{c} \overline{\nu}_{e} \rightarrow \overline{\nu}_{e} \\ (\nu_{e} \rightarrow \nu_{e}) \end{array} $	Baksan, LENS, Borexino, SNO+, Richochet, CeLAND, Daya-Bay
Isotope Source	Disapp	$\stackrel{-}{v_{e}} \rightarrow \stackrel{-}{v_{e}}$	IsoDAR
Pion / Kaon Decay-at-Rest Source	Appearance & Disapp	$ \begin{array}{c} \overline{\nu}_{\mu} \rightarrow \overline{\nu}_{e} \\ \nu_{e} \rightarrow \nu_{e} \end{array} $	OscSNS, CLEAR, DAEδALUS, KDAR
Accelerator v / $\stackrel{-}{v}$ using Pion Decay-in-Flight	Appearance & Disapp	$\begin{array}{c} v_{\mu} \rightarrow v_{e} \;,\; \stackrel{-}{v_{\mu}} \rightarrow \stackrel{-}{v_{e}} \\ v_{\mu} \rightarrow v_{\mu} \;,\; v_{e} \rightarrow \; v_{e} \end{array}$	MINOS+, MicroBooNE, LAr1kton+MicroBooNE, CERN SPS
Low-Energy ν-Factory	Appearance & Disapp	$\begin{array}{c} \nu_{e} \rightarrow \nu_{\mu} \; , \; \stackrel{-}{\nu_{e}} \rightarrow \stackrel{-}{\nu_{\mu}} \\ \nu_{\mu} \rightarrow \nu_{\mu} \; , \; \nu_{e} \rightarrow \; \nu_{e} \end{array}$	vSTORM at Fermilab

# **Very Short Baseline Exps and Project-X**

- See talk by Heather Ray
- Some comments and questions:
  - Need large detectors (>1 kton) with capability to detect IBD events
    - Best to see oscillations within the detector
    - Need PMT coverage to be able to see neutron capture on hydrogen
  - Neutrino sources
    - Isotope source using 60 to 100 MeV with 200 kW to 600 kW
    - DAR sources using 800 MeV proton beams of 10 kW to 100 kW
  - Can one use timing to overcome backgrounds rather that going deep underground?
    - Isotope source Probably not.
      - 8Li halflife 840 msec
    - DAR Source May be possible
      - Neutron capture takes 190 μsec on hydrogen
      - But need big enough detector to see oscillatory behavior

# **Backup**

# MiniBooNE $v_{\mu}$ and $\overline{v}_{\mu}$ Disappearance Limits



# IsoDAR sin<sup>2</sup>θ<sub>W</sub> Measurement

- Weak mixing (Weinberg) angle  $\theta_W$ 
  - Measured very precisely by LEP experiments
  - NuTeV neutrino-quark scattering measurement ~3σ high (NuTeV anomaly)
- Measure in IsoDAR using  $\overline{v}_e$  +  $e^- \rightarrow \overline{v}_e$  +  $e^-$ 
  - If IsoDAR also sees discrepancy then this could be new physics associated with neutrinos
  - If IsoDAR does not see a discrepancy then NuTeV Anomaly something to do with quark distributions or other quark physics.

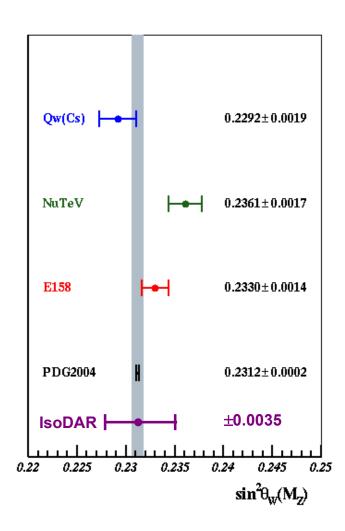
Similar to: PHYSICAL REVIEW D 71, 073013 (2005)

#### Precision measurement of $\sin^2 \theta_W$ at a reactor

J. M. Conrad, J. M. Link, and M. H. Shaevitz

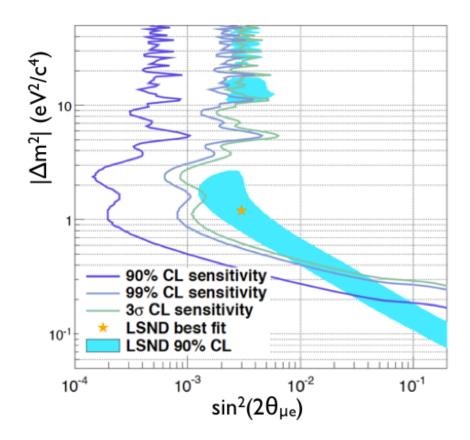
Department of Physics, Columbia University, New York, New York 10027, USA

(Received 22 July 2004; published 28 April 2005)



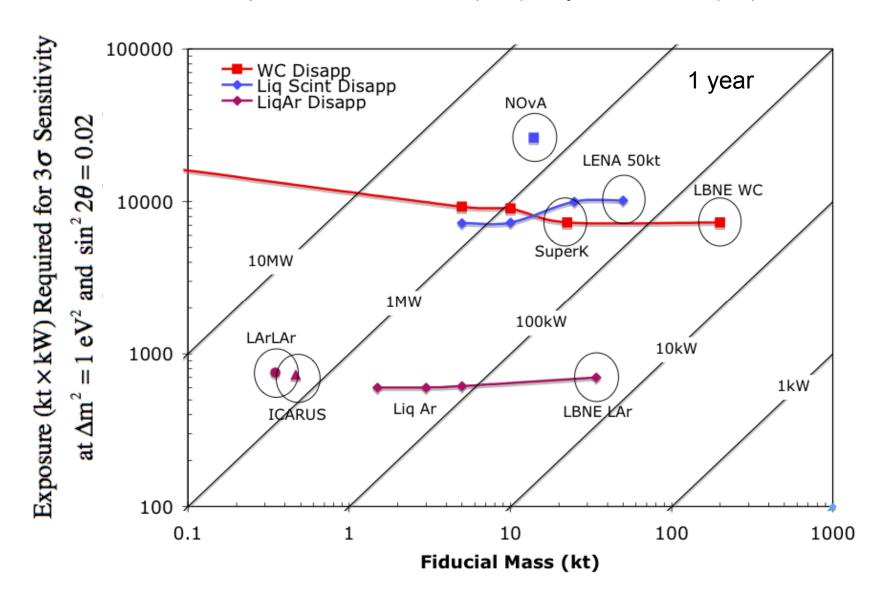
# Using Coherent v-Nucleon Scattering with DAR Source

- Coherent process sensitive to all active neutrinos so any disappearance would indicate osc to sterile neutrinos
- High cross section but very low recoil energy (few to tens of keV)
  - Need to use high sensitivity detectors: LAr (LNe),CDMS,COUPP
- CLEAR proposal: 450kg LAr 46m from dump at SNS (arXiv:0910.1989)
- DAEδALUS DAR source: 100kg Ge detector at 10m (>2000 evnts/yr)



# VSBL $v_e$ Disappearance: Source Power and Detector Size for x10 Better Sensitivity Than Current

Exposure = Detector Size (kton) × Cyclotron Power (kW)



# Example (3+1) and (3+2) Model Fits

### 3+1 Model:

$$P(\nu_{\mu} \to \nu_{e}) = 4|U_{e4}|^{2}|U_{\mu4}|^{2}\sin^{2}x_{41} \qquad P(\nu_{e} \to \nu_{e}) = 1 - 4|U_{e4}|^{2}(1 - |U_{e4}|^{2})\sin^{2}x_{41}$$
$$= \sin^{2}2\theta_{\mu e}\sin^{2}x_{41} \qquad = 1 - \sin^{2}2\theta_{ee}\sin^{2}x_{41}$$

Example Fit:  $\Delta m_{41}^2 = 0.92 \, eV^2$   $\sin^2 2\theta_{\mu e} = 0.0025$   $\sin^2 2\theta_{\mu \mu} = 0.13$   $\sin^2 2\theta_{ee} = 0.073$ 

G. Karagiorgi, Z. Djurcic, J. Conrad, M. Shaevitz, and M. Sorel,

Phys.Rev. D80, 073001 (2009), 0906.1997

$$x_{ij} \equiv \mathring{\Delta} m_{ij}^2 L/4E$$

#### 3+2 Model:

$$P(\nu_{\mu} \to \nu_{e}) = 4|U_{e4}|^{2}|U_{\mu4}|^{2}\sin^{2}x_{41} + P(\nu_{\alpha} \to \nu_{\alpha}) = 1 - 4[(1 - |U_{\alpha4}|^{2} - |U_{\alpha5}|^{2}) \cdot 4|U_{e5}|^{2}|U_{\mu5}|^{2}\sin^{2}x_{51} + (|U_{\alpha4}|^{2}\sin^{2}x_{41} + |U_{\alpha5}|^{2}\sin^{2}x_{51}) + 8|U_{e4}U_{\mu4}U_{e5}U_{\mu5}|\sin x_{41}\sin x_{51}\cos(x_{54} \pm \delta) + |U_{\alpha4}|^{2}|U_{\alpha5}|^{2}\sin^{2}x_{54}]$$

J. Kopp, M. Maltoni, and

T. Schwetz (2011), 1103.4570.

(Short baseline approximation where highest mass state dominates:  $\Delta m_{12}^2 \approx \Delta m_{13}^2 \approx 0$ )